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# STUDY OF STERILE NEUTRINO CONTRIBUTION TO NEUTRINOLESS DOUBLE BETA DECAY

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#### Abstract

The sterile neutrino contribution to neutrinoless double beta  $(0\nu\beta\beta)$  decay has been first studied within a specific model with the two sterile neutrinos mixed only with the  $v_e$  flavor state. In the present paper, we assume that in addition to the three conventional light neutrinos there exists only one Majorana neutrino mass-eigenstate  $v_h$ , dominated by the sterile neutrino species, with an arbitrary mass  $m_h$ , which may mix with all the active neutrino weak eigenstates  $v_{e,\mu,\tau}$ . We study possible contribution of this  $v_h$  neutrino state to  $0\nu\beta\beta$ -decay via a nonzero admixture of  $v_e$  weak eigenstate.

**Keywords:** Sterile neutrino, double beta decay, neutrino mass, PHFB model, nuclear matrix elements

# 1. Introduction

The sterile right-handed neutrinos not participating in the electroweak interactions, are natural candidates for the extension of the standard model (SM) [1]. It has long been recognized that the

sterile neutrinos may have plenty of phenomenological implications. One of them is the seesaw generation of tiny Majorana neutrino masses for the three observable very light neutrinos via a very large Majorana mass term of the right-handed neutrinos. In this case together with the very light neutrinos there also appear very heavy Majorana neutrino mass states with the typical masses of  $10^{12}$  GeV. However, in a more general case of the sterile-active neutrino mixing there may also appear additional neutrino states  $v_h$  with arbitrary masses  $m_h$ . The neutrino mass eigenstates  $v_h$ , dominated by the sterile  $v_R$  neutrino weak eigenstates, contain some admixture of the active  $v_{e,\mu,\tau}$  neutrino species that allows the  $v_h$  to contribute to various processes, in particular, to those which are forbidden in the SM by the Lepton Number or Lepton Flavor Violation (LNV or LFV). They may also modify the interpretation of cosmological and astrophysical observations. Therefore, the masses  $m_h$  of  $v_h$  neutrino states and their mixing with the active neutrinos are subject to various experimental as well as cosmological and astrophysical constraints. Various implications of the sterile neutrinos have been extensively studied in the literature [2-7].

The sterile neutrino contribution to  $0\nu\beta\beta$ -decay has been first studied in the pioneer work [8] within a specific model with the two sterile neutrinos mixed only with the  $v_e$  flavor state. In the present case, we assume that in addition to the three conventional light neutrinos there exists only one Majorana neutrino mass eigen state  $v_h$ , dominated by the sterile neutrino species, with an arbitrary mass  $m_h$ , which may mix with all the active neutrino weak eigenstates,  $v_{e,\mu,\tau}$ . We study possible contribution of this  $v_h$  neutrino state to  $0\nu\beta\beta$ -decay via a nonzero admixture of  $v_e$  weak eigenstate. The paper is organized as follows. In Section 2, theoretical description of the mechanism and model used to calculate matrix elements is presented. Section 3 deals with results and discussions. Some Concluding remarks are presented in Section 4.

# 2. Mechanism involving sterile neutrinos

The contribution of the sterile  $v_h$  neutrino to the half-life  $T_{1/2}^{0\nu}$  for the  $0^+ \rightarrow 0^+$  transition of  $\partial v\beta\beta$ -decay has been derived by considering the exchange of a Majorana neutrino between two nucleons and is given by [9]

$$\left[T_{1/2}^{0\nu}(0^+ \to 0^+)\right]^{-1} = G_{01} \left| U_{eh}^2 \frac{m_h}{m_e} M^{(\nu)}(m_h) \right|^2$$
(1)

Here  $m_e$  is the electron mass and  $G_{01}$  is the phase space factor given by

$$G_{01} = \left[\frac{2(G_F g_A)^4 m_e^9}{64\pi^5 (m_e R)^2 \ln (2)}\right] \times \int_{1}^{T+1} F_0(Z_f, \varepsilon_1) F_0(Z_f, \varepsilon_2) p_1 p_2 \varepsilon_1 \varepsilon_2 d\varepsilon_1$$
(2)

 $U_{eh}$  is the  $v_h - v_e$  mixing matrix element and the NTME  $M^{(0v)}(m_h)$  is written as

$$M^{(0\nu)}(m_h) = \langle 0_F^+ \left\| \left[ -\frac{H_F(m_h, r)}{g_A^2} + \sigma_n \cdot \sigma_m H_{GT}(m_h, r) + S_{nm} H_T(m_h, r) \right] \tau_n^+ \tau_m^+ \right\| 0_I^+ \rangle$$
(3)

with  $S_{nm} = 3(\sigma_n, \mathbf{r}_{nm})(\sigma_m, \mathbf{r}_{nm}) - \sigma_n, \sigma_m$ 

The neutrino potentials  $H_{\alpha}$  are of the form

$$H_{\alpha}(m_{h},r) = \frac{2R}{\pi} \int_{0}^{\infty} \frac{f_{\alpha}(qr)h_{\alpha}(q^{2})q^{2}dq}{\sqrt{q^{2} + m_{h}^{2}}\left(\sqrt{q^{2} + m_{h}^{2}} + \overline{A}\right)}$$
(4)

where  $f_{\alpha}(qr) = j_0(qr)$  and  $f_{\alpha}(qr) = j_2(qr)$  for  $\alpha$ = Fermi (F)/Gamow Teller (GT) and tensor (T) potentials respectively. The effects due to the finite size (FNS) are incorporated through the dipole form factors [10]. The functions  $h_F(q)$ ,  $h_{GT}(q)$  and  $h_T(q)$  are given by

$$h_{F}(q) = g_{\nu}^{2}(q^{2})$$
(5)  

$$h_{GT}(q) = \frac{g_{A}^{2}(q^{2})}{g_{A}^{2}} \left[ 1 - \frac{2}{3} \frac{g_{P}(q^{2})q^{2}}{g_{A}(q^{2})2M_{P}} + \frac{1}{3} \frac{g_{P}^{2}(q^{2})q^{4}}{g_{A}^{2}(q^{2})4M_{P}^{2}} \right] + \frac{2}{3} \frac{g_{M}^{2}(q^{2})q^{2}}{g_{A}^{2}4M_{P}^{2}}$$
$$\approx \left( \frac{\Lambda_{A}^{2}}{q^{2} + \Lambda_{A}^{2}} \right)^{4} \left[ 1 - \frac{2}{3} \frac{q^{2}}{(q^{2} + m_{\pi}^{2})} + \frac{1}{3} \frac{q^{4}}{(q^{2} + m_{\pi}^{2})^{2}} \right]$$
$$+ \left( \frac{g_{V}}{g_{A}} \right)^{2} \frac{K^{2}q^{2}}{6M_{P}^{2}} \left( \frac{\Lambda_{V}^{2}}{q^{2} + \Lambda_{V}^{2}} \right)^{4}$$
(6)  

$$h_{T}(q) = \frac{g_{A}^{2}(q^{2})}{g_{A}^{2}} \left[ \frac{2}{3} \frac{g_{P}(q^{2})q^{2}}{g_{A}(q^{2})2M_{P}} - \frac{1}{3} \frac{g_{P}^{2}(q^{2})q^{4}}{g_{A}^{2}(q^{2})4M_{P}^{2}} \right] + \frac{1}{3} \frac{g_{M}^{2}(q^{2})q^{2}}{g_{A}^{2}4M_{P}^{2}},$$
$$\approx \left( \frac{\Lambda_{A}^{2}}{q^{2} + \Lambda_{A}^{2}} \right)^{4} \left[ \frac{2}{3} \frac{q^{2}}{(q^{2} + m_{\pi}^{2})} - \frac{1}{3} \frac{q^{4}}{(q^{2} + m_{\pi}^{2})^{2}} \right]$$
$$+ \left( \frac{g_{V}}{g_{A}} \right)^{2} \frac{K^{2}q^{2}}{12M_{P}^{2}} \left( \frac{\Lambda_{V}^{2}}{q^{2} + \Lambda_{V}^{2}} \right)^{4}$$
(7)

where

$$g_V(q^2) = g_V \left(\frac{\Lambda_V^2}{q^2 + \Lambda_V^2}\right)^2 \tag{8}$$

$$g_A(q^2) = g_A \left(\frac{\Lambda_A^2}{q^2 + \Lambda_A^2}\right)^2 \tag{9}$$

$$g_P(q^2) = \frac{2M_P g_A(q^2)}{(q^2 + m_\pi^2)} \left(\frac{\Lambda_A^2 - m_\pi^2}{\Lambda_A^2}\right)$$
(10)  
$$g_M(q^2) = K g_V(q^2)$$
(11)

with  $g_V = 1.0, g_A = 1.254, K = \mu_p - \mu_n = 3.70$ ,  $\Lambda_V = 0.850 GeV$ , and  $\Lambda_A = 1.086 GeV$ .

The short range correlation (SRC) has been included using the Jastrow correlation with Miller Spencer type of parametrization given by [11],

$$f(r) = 1 - ce^{-ar^2}(1 - br^2)$$
(12)

with  $a = 1.1 \text{ fm}^{-2}$ ,  $b = 0.68 \text{ fm}^{-2}$  and c = 1.0. The above matrix elements are calculated with the help of Projected Hartree Fock Bogoliubov (PHFB) model.

#### 2.1 PHFB model

The expression to calculate the NTMEs  $M_{\alpha}$  of  $(\beta^{-}\beta^{-})_{0\nu}$  decay for the  $0^{+} \rightarrow 0^{+}$  transition in the PHFB model is obtained as follows. In the PHFB model, a state with good angular momentum J is obtained from the axially symmetric HFB intrinsic state  $|\phi_{0}\rangle$  with K = 0 through the following relation using the standard projection technique [12]:

$$\left|\psi_{00}^{J}\right\rangle = \left[\frac{(2J+1)}{8\pi^{2}}\right] \int D_{00}^{J}(\Omega)R(\Omega)\left|\phi_{0}\right\rangle d\Omega \qquad (13)$$

Where  $R(\Omega)$  and  $D^{J}_{MK}(\Omega)$  are the rotation operator and the rotation matrix, respectively. The axially symmetric HFB intrinsic state  $|\phi_0\rangle$  can be written as

$$|\phi_0\rangle = \prod_{im} \left( u_{im} + v_{im} b^{\dagger}_{im} b^{\dagger}_{i\widetilde{m}} \right) |0\rangle$$
(14)

Where the creation operators  $b_{im}^{\dagger}$  and  $b_{i\tilde{m}}^{\dagger}$  are defined as

$$b_{im}^{\dagger} = \sum_{\alpha} C_{i\alpha,m} a_{\alpha m}^{\dagger}$$
(15)

And

$$b_{i\tilde{m}}^{\dagger} = \sum_{\alpha} (-1)^{l+j-m} C_{i\alpha,m} a_{\alpha,-m}^{\dagger}$$
(16)

The amplitude  $(u_{im}, v_{im})$  and the expansion coefficient  $C_{ij,m}$  are obtained from the HFB calculations. Employing the HFB wave functions, one obtains the following expression for NTMEs  $M_{\alpha}$  of  $(\beta^{-}\beta^{-})_{0\nu}$  decay [13,14]

$$\begin{split} \mathsf{M}\alpha &== [\mathsf{n}^{\mathsf{J}_{\mathsf{i}}=0}\mathsf{n}^{\mathsf{J}_{\mathsf{f}}=0}]^{-1/2} \\ \times \int_{0}^{\pi} n_{(Z,N),(Z+2,\,N-2)}(\theta) \frac{1}{4} \sum_{\alpha\beta\gamma\delta} \langle \alpha\beta | \mathcal{O}_{\alpha} | \gamma\delta \rangle \times \sum_{\varepsilon n} \frac{\left(f_{Z+2,N-2}^{(\pi)*}\right)_{\varepsilon\beta}}{\left[1 + F_{Z,N}^{(\pi)}(\theta) f_{Z+2,N-2}^{(\pi)*}\right]_{\varepsilon\alpha}^{-1}} \\ \times \frac{\left(f_{Z,N}^{(\nu)*}\right)_{\eta\delta}}{\left[1 + F_{Z,N}^{(\nu)}(\theta) f_{Z+2,N-2}^{(\nu)*}\right]_{\gamma\eta}^{-1}} \sin\theta \, d\theta \end{split}$$
(17)

where

$$n^{J} = \int_{0}^{\pi} \left[ det \left( 1 + F^{(\pi)} f^{\pi^{\dagger}} \right) \right]^{1/2} \times \left[ det \left( + F^{(\nu)} f^{\nu^{\dagger}} \right) \right]^{1/2} d_{00}^{J}(\theta) \sin(\theta) d\theta$$
(18)

$$n_{(Z,N),(Z+2,N-2)}(\theta) = \left[\det\left(1 + F_{Z,N}^{(\nu)}f_{Z+2,N-2}^{(\nu)^{\dagger}}\right)\right]^{1/2} \times \left[\det\left(1 + F_{Z,N}^{(\pi)}f_{Z+2,N-2}^{(\pi)^{\dagger}}\right)\right]^{1/2}$$
(19)

The  $\pi(v)$  represents the proton (neutron) of nuclei involved in the  $(\beta\beta)_{\theta v}$  decay process. The matrix  $F_{Z,N}(\theta)$  and  $f_{Z,N}$  are given by

$$F_{Z,N}(\theta) = \sum_{m'_{\alpha},m'_{\beta}} d^{j_{\alpha}}_{m_{\alpha},m'_{\alpha}}(\theta) d^{j_{\beta}}_{m_{\beta},m'_{\beta}}(\theta) f_{j_{\alpha}m'_{\alpha},j_{\beta}m'_{\beta}}(20)$$
$$f_{Z,N} = \sum_{i} C_{ij_{\alpha},m_{\alpha}} C_{ij_{\beta},m_{\beta}} (v_{im_{\alpha}}/u_{im_{\alpha}}) \delta_{m_{\alpha},-m_{\beta}} (21)$$

The calculations of required NTMEs  $M_{\alpha}$  are performed in the following manner. In the first step, matrices  $F_{Z,N}(\theta)$  and  $f_{Z,N}$  are set up for the nuclei involved in the  $(\beta\beta)_{0\nu}$  decay making use of 20 Gaussian quadrature points in the range  $(0,\pi)$ . Finally, the required NTMEs can be calculated in a straightforward manner using Eqn. (21).

### 3. **Results and discussions**

We have calculated the NTMEs defined in Eqn. (3) and results are shown in table 1. We have also calculated the expected half-life  $T_{1/2}^{0\nu}$  for a single neutrino of mass m<sub>h</sub> with coupling  $U_h$ , shown in table 2. These half-lives are calculated for axial vector coupling constant g<sub>A</sub>=1.254 and plotted in fig. 1-3. In the same figures. current limits from the experiment AMoRE-I for <sup>100</sup>Mo [15], for <sup>128</sup>Te [16] and from CUORE for <sup>130</sup>Te [17] are also shown. The variation of these half-lives with the coupling  $U_{eh}^2$  is plotted in fig. 1-3 for sterile neutrino mass ranging from 100 eV to 10 GeV. It is observed that

half-life increases for lower values of mass  $m_N$  in all three cases. Also the half-lives increase when  $U_{eh}^2$  changes from  $10^{-2}$  to  $10^{-8}$ .

Table 1: Calculated NTMEs  $M^{(\nu)}(m_h)$  for  $\theta \nu \beta \beta$ -decay of <sup>100</sup>Mo and <sup>128,130</sup>Te isotope including short range correlation (SRC)

Nuclei	Mass $(m_h)$ (GeV)						
	0.001	0.01	0.1	1.0	10	100	
<sup>100</sup> Mo	6.262	5.91	3.033	0.105	0.998	0.998	
<sup>128</sup> Te	3.619	3.412	1.64	0.064	0.613	0.613	
<sup>130</sup> Te	4.054	3.815	1.808	0.069	0.659	0.659	

Table 2: Half life  $T_{1/2}^{0\nu}$  for a single neutrino of mass m<sub>h</sub> with coupling  $U_{eh}^2$  from 10<sup>-2</sup> to 10<sup>-8</sup> for <sup>100</sup>Mo, <sup>128,130</sup>Te isotopes

Nuclei	Mass (eV)	$T_{1/2}^{0\nu}$ (years)					
		$U_{eh}^2 = 10^{-2}$	10-4	10-6	10-8		
<sup>100</sup> Mo	10 <sup>2</sup>	1.43x10 <sup>23</sup>	1.43x10 <sup>27</sup>	1.43x10 <sup>31</sup>	1.43x10 <sup>35</sup>		
	104	1.43x10 <sup>19</sup>	$1.43 \times 10^{23}$	$1.43 \times 10^{27}$	$1.43 \times 10^{31}$		
	106	1.43x10 <sup>15</sup>	1.43x10 <sup>19</sup>	$1.43 \times 10^{23}$	$1.43 \times 10^{27}$		
	108	6.12x10 <sup>11</sup>	$6.12 \times 10^{15}$	6.12x10 <sup>19</sup>	$6.12 \times 10^{23}$		
	10 <sup>10</sup>	5.65x10 <sup>8</sup>	5.65x10 <sup>12</sup>	5.65x10 <sup>16</sup>	$5.65 \times 10^{20}$		
<sup>128</sup> Te	10 <sup>2</sup>	1.08x10 <sup>25</sup>	1.08x10 <sup>29</sup>	1.08x10 <sup>33</sup>	1.08x10 <sup>37</sup>		
	104	$1.08 \times 10^{21}$	$1.08 \times 10^{25}$	1.08x10 <sup>29</sup>	$1.08 \times 10^{33}$		
	10 <sup>6</sup>	$1.08 \mathrm{x} 10^{17}$	$1.08 \times 10^{21}$	$1.08 \times 10^{25}$	$1.08 \times 10^{29}$		
	10 <sup>8</sup>	5.25x10 <sup>13</sup>	5.25x10 <sup>17</sup>	$5.25 \times 10^{21}$	$5.25 \times 10^{25}$		
	1010	3.76x10 <sup>10</sup>	3.76x10 <sup>14</sup>	3.76x10 <sup>18</sup>	3.76x10 <sup>22</sup>		
<sup>130</sup> Te	10 <sup>2</sup>	3.53x10 <sup>23</sup>	$3.53 \times 10^{27}$	3.53x10 <sup>31</sup>	3.53x10 <sup>35</sup>		
	10 <sup>4</sup>	3.53x10 <sup>19</sup>	$3.53 \times 10^{23}$	3.53x10 <sup>27</sup>	3.53x10 <sup>31</sup>		

106	$3.53 \times 10^{15}$	3.53x10 <sup>19</sup>	$3.53 \times 10^{23}$	$3.53 \times 10^{27}$
108	$1.78 \times 10^{12}$	$1.78 \times 10^{16}$	$1.78 \times 10^{20}$	1.78x10 <sup>24</sup>
10 <sup>10</sup>	1.34x10 <sup>9</sup>	1.34x10 <sup>13</sup>	1.34x10 <sup>17</sup>	1.34x10 <sup>21</sup>

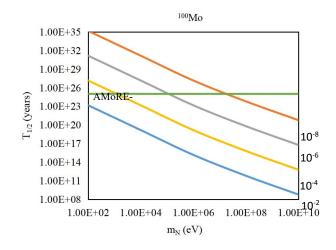


Figure 1: Half-life for a single neutrino of mass  $m_N$  with coupling  $U_{eh}^2 = 10^{-2}$  to  $10^{-8}$ . Green line represents the AMoRE-I experiment limit for <sup>100</sup>Mo

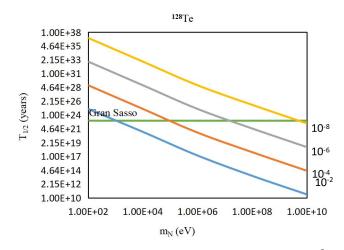
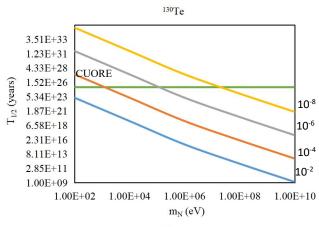


Figure. 2: Half-life for a single neutrino of mass  $m_N$  with coupling  $U_{eh}^2 = 10^{-2}$  to  $10^{-8}$ . Green line represents the Gran Sasso experiment limit for <sup>128</sup>Te



# Figure. 3: Half-life for a single neutrino of mass $m_N$ with coupling $U_{eh}^2 = 10^{-2}$ to $10^{-8}$ . Green line represents the CUORE experiment limit for <sup>130</sup>Te

The analysis of light sterile neutrino of eV order mass can be done by writing [18]

$$\langle m_{light} \rangle = \sum_{k=1}^{3} U_{ek}^{2} m_{k} + \sum_{i} U_{ei}^{2} m_{i}$$
 (22)

where k stands for three known neutrino species and I for unknown neutrinos of eV order mass and comparing the experimental limits with the calculated half-lives using

$$\left[T_{1/2}^{0\nu}(0^+ \to 0^+)\right]^{-1} = G_{01} \left[\frac{\langle m_{light} \rangle}{m_e}\right]^2 |M^\nu|^2$$
(23)

we get the results shown in table 3. We have presented the extracted limits on effective neutrino mass  $\langle m_{light} \rangle$  by using the phase space factors G<sub>01</sub> from [23] and available experimental half-lives for considered isotopes. It is observed that best upper limit of 0.22 eV is obtained in case of <sup>130</sup>Te isotope. Also we have shown the predicted half-lives for considered transitions taking  $\langle m_{light} \rangle = 50$  meV and shown the results in column 7 of the table 3. The half-lives are of the order of  $10^{26}$  years as observed from the above table.

Table 3: Experimental half-lives along with extracted limits on effective neutrino mass and predicted half-lives for  $\langle m_{light} \rangle = 50 \text{meV}$ 

Decay	Μ <sup>ν</sup>	$T^{0\nu}_{1/2}$	Ref.	G <sub>01</sub> (yr	$\langle m_{light} \rangle$	$T_{1/2}^{0\nu}$ (yrs)
		(yrs)		<sup>1</sup> )	(eV)	(predicted)
		(Exp.)				
<sup>96</sup> Zr→ <sup>96</sup> Mo	2.624	9.2x10 <sup>21</sup>	[19]	2.06x10 <sup>-</sup> 14	14.15	7.36x10 <sup>26</sup>
<sup>100</sup> M0→ <sup>100</sup> Ru	5.871	1.1x10 <sup>24</sup>	[20]	1.59x10 <sup>-</sup> 14	0.66	1.90x10 <sup>26</sup>
<sup>110</sup> Pd→ <sup>110</sup> Cd	6.977	6.0x10 <sup>17</sup>	[21]	4.83x10 <sup>-</sup>	1.36x10 <sup>3</sup>	4.44x10 <sup>25</sup>
<sup>128</sup> Te→ <sup>128</sup> Xe	2.906	1.1x10 <sup>23</sup>	[16]	1.66x10 <sup>-</sup>	13.01	7.44x10 <sup>27</sup>
<sup>130</sup> Te→ <sup>130</sup> Xe	4.029	2.3x10 <sup>25</sup>	[17]	1.43x10 <sup>-</sup>	0.22	4.49x10 <sup>26</sup>

 $^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$  2.931 2.0x10<sup>22</sup> [22] 6.32x10<sup>-</sup> 1.55 1.92x10<sup>26</sup> 14

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# 4. Conclusions

We have calculated the NTMEs of  $0\nu\beta\beta$  decay process associated with light sterile neutrino in PHFB model and using Eqn. (5), we have obtained the half-life  $T_{1/2}^{0\nu}$  for a single neutrino of mass  $m_h$  with  $v_h - v_e$  mixing matrix element coupling  $U_{eh}^2$  from 10<sup>-2</sup> to 10<sup>-8</sup> for <sup>100</sup>Mo, <sup>128,130</sup>Te isotopes. We have also shown the effect of  $U_{eh}^2$  on half-lives of these isotopes in fig. 1-3. It is observed that half-life increases for lower values of mass  $m_N$  in all three cases. Also the half-lives increase when  $U_{eh}^2$  changes from 10<sup>-2</sup> to 10<sup>-8</sup>. The predicted half-lives are of the order of 10<sup>26</sup> years as shown in table 3 for effective light neutrino mass of 50 meV.

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